## Multi-View 3D Reconstruction of Highly-Specular Objects

**Master Thesis** 



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## **Motivation**

- Goal: faithful reconstruction of full 3D shape of an object
- Current techniques:















## **Motivation**

- Challenge: objects exhibiting a complex reflectance behavior
- Focus: on opaque+specular materials







### Introduction

- **Observation:** shading is a powerful visual cue
- Provides surface orientation information







## Introduction

 Many techniques for surface normal field estimation using shading cues from single view



- How can information from several viewpoints be combined?
- Could 3D reconstruction of specular objects be addressed this way?



Image credits: N. Funk and Y.-H. Yang. Using a Raster Display Device for Photometric Stereo.



## Background

#### Mesostructure from Specularity (Chen et al., CVPR '06)



 Gloss and Normal Map Acquisition of Mesostructures Using Gray Codes (Francken et al., ISVC '09)





## Background

- Specularity-Consistency based methods
  - Voxel Carving for Specular Surfaces (Bonfort et al., ICCV '03)
  - Dense 3D Reconstruction from Specularity Consistency (Nehab et al., CVPR '08)



Image credits: T. Bonfort and P. Sturm. Voxel carving for specular surfaces *(left)*, J. Balzer and S.Werling. Principles of Shape from Specular Reflection *(right)* 

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## Background

Results







(a) Captured normals

(e) Captured normals



(b) Captured positions

(f) Captured positions

(c) Refined positions





(g) Refined positions



(d) Traditional scanner



(h) Photograph





#### **Our Approach**

- Estimate multi-view normal fields using structured environment
- Multi-view normal field integration problem
  - Need very robust algorithm!



## **Multi-View Normal Field Integration**

Synthetic normals

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#### Chang et al., CVPR'07

• Level sets





1

- Master thesis of Z. Dai
  - MRF approach



### **Overview**

#### I. Multi-View Normal Field Integration

#### II. Multi-View Shape-from-Specularity

#### III. Evaluation

#### **IV.** Conclusion and Future Work





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## **Problem Statement**

### **Given**:

- $\kappa_c$  calibrated cameras  $C_i$
- Projection matrices  $P_i = K_i [R_i | t_i] \mid i = 1 \dots \kappa_c$
- Normal fields  $\mathcal{N}_i$

### Goal:

• Reconstruction of surface  $\partial S$ 

### Problem:

 Inferring coordinates of all surface point given normal fields estimates









## Challenges

- Noise
- Outliers
- Systematic errors
- Holes
- Initial guess (visual hull) in practice difficult to compute
- Implementation concerns
  - Fine-detail reconstruction



## **Variational Approach**

- Solve variational problem:  $E(\partial S) = \lambda_1 \underbrace{\int_{\partial S} dA - \lambda_2}_{E_1} \underbrace{\int_{\partial S} \langle c\mathbf{N}, \mathbf{n} \rangle \ dA}_{E_2},$ 
  - **N**(**x**) ... vector field, reconstructed from normal fields
  - c(**x**) ... surface consistency
- Solving minimal surface problems
  - Active contour
  - Level sets
  - Graph cuts
  - Convex relaxation







## **Vector Field Computation: Idea**

- Core of our approach
- Key questions:
  - Surface consistency measure  $c(\mathbf{x}) : \mathbf{x} \mapsto \mathbb{R}$
  - Value of vector field  $\mathbf{N}(\mathbf{x}): \mathbf{x} \mapsto \mathbb{S}^2$



V

 $\mathbf{x} \in V \subset \mathbb{R}^3$ 



## **Vector Field Computation**

- Back-project normal fields into volume V
- Map back-projected normals to feature space
- Density estimation to find the patterns normal
  - Based on discrete, back-projected normal samples
  - Non-parametric method essential



## **Feature Space Analysis**

- Histogram method
- Kernel density estimation
- Mean-Shift clustering





Image credits: Christopher M. Bishop. Pattern Recognition and Machine Learning (Information Science and Statistics) *(left)*, D. Comaniciu and P. Meer. Mean shift: A robust approach toward feature space analysis *(right)*, Universitätbonn

## Implementation

- Octree-based discretization of bounding volume V
- Initial refinement strategy
- Continuous Max-Flow based volume segmentation
- Iterative scheme



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#### I. Multi-View Normal Field Integration

#### II. Multi-View Shape-from-Specularity

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## **Multi-View Shape-from-Specularity**

How to compute normal fields of specular objects?



#### Challenges

- How to lit object fully?
- Distant light assumption violation
- How to reliably decode patterns?
- Screen calibration



### **Proposed Setup**



## **Capturing the Data**







## **Computing Light Maps**

Fuzzy decoding









### **Screen Calibration**

• Structured pattern based triangulation  $Q = \sum_{m=1}^{M} \left( \mathbf{p}_m - (\mathbf{o} + q_{x,m}\mathbf{a} + q_{y,m}\mathbf{b}) \right)^2$ 



![](_page_24_Picture_3.jpeg)

### **Normal-Depth Ambiguity**

![](_page_25_Figure_1.jpeg)

### Reconstruction

- Input are light-maps and not normal fields
- Project labels to octree corners and compute normal hypotheses
- Reconstruct vector field and compute surface consistency
- Fit surface to vector field

![](_page_26_Picture_5.jpeg)

![](_page_26_Picture_6.jpeg)

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![](_page_27_Picture_4.jpeg)

![](_page_27_Picture_5.jpeg)

### **Synthetic Normal Fields**

![](_page_28_Picture_1.jpeg)

### **Synthetic Normal Fields**

![](_page_29_Picture_1.jpeg)

## **Real Data: Photometric Stereo**

- Classic least-squares photometric stereo
  - Simple thresholding prior to fitting
- 6 cameras, 12 rotations, 72 views
- 198 images for computation of single normal field

![](_page_30_Picture_5.jpeg)

![](_page_30_Picture_6.jpeg)

![](_page_30_Picture_7.jpeg)

### **Real Data: Photometric Stereo**

![](_page_31_Picture_1.jpeg)

![](_page_31_Picture_2.jpeg)

- 10 cameras, 24 rotations, 240 views, two sources of structured illumination
- Mirror bunny

![](_page_32_Picture_3.jpeg)

![](_page_32_Picture_4.jpeg)

![](_page_32_Picture_5.jpeg)

![](_page_32_Picture_6.jpeg)

![](_page_33_Picture_1.jpeg)

![](_page_34_Picture_1.jpeg)

![](_page_35_Picture_1.jpeg)

![](_page_35_Picture_2.jpeg)

#### I. Multi-View Normal Field Integration

#### II. Multi-View Shape-from-Specularity

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![](_page_36_Picture_4.jpeg)

![](_page_36_Picture_5.jpeg)

## Conclusion

- New, robust multi-view normal field integration algorithm
  - No initial guess (visual hull) needed
  - First results, demonstrated on captured data
  - Efficient numerical techniques
- New dome-based method for reconstruction of highly specular objects
  - Display screens as sources of structured lighting
  - Normal computation and integration based approach
  - State-of-the-art results

![](_page_37_Picture_9.jpeg)

### **Future Work**

- Testing integration algorithm using more general normal estimation techniques
- Coding the light pattern using different coding strategies
- Weight normals coming from specular surfaces according to uncertainty of source of illumination
- Parallelization potential

![](_page_38_Picture_5.jpeg)

![](_page_38_Picture_6.jpeg)

# Thank you for your attention!

![](_page_39_Picture_1.jpeg)

![](_page_39_Picture_2.jpeg)

## **Structured Light 3D Scanning**

![](_page_40_Figure_1.jpeg)

![](_page_40_Figure_2.jpeg)

![](_page_40_Picture_3.jpeg)

Image credits: J. Geng: Structured-Light 3D Surface Imaging: A Tutorial

![](_page_40_Picture_5.jpeg)

### **Photometric Stereo**

Lambertian assumption:

![](_page_41_Picture_3.jpeg)

![](_page_41_Picture_4.jpeg)

Image credits: R. Basri, D. Jacobs, I. Kemelmacher: Photometric Stereo with General, Unknown Lighting (*bottom image*)

![](_page_41_Picture_6.jpeg)

### **Convex Relaxation**

$$E(\gamma) = \lambda_1 \int_V \|\nabla \gamma\| \ dV - \lambda_2 \int_V (\nabla \cdot (c\mathbf{N})) \gamma \ dV$$

 $w.r.t \ \gamma: V \mapsto [0,1]$ 

![](_page_42_Picture_3.jpeg)

![](_page_42_Picture_4.jpeg)

### **Reflectance Models**

![](_page_43_Figure_1.jpeg)

![](_page_43_Picture_2.jpeg)

Image credits: H. T. Nefs, J. J. Koenderink, A. M.L. Kappers: Shape-from-Shading for Matte and Glossy Objects (*bottom*).

![](_page_43_Picture_4.jpeg)